

#### **Philosophical Transactions**

Please note: Due to an error in the print volume, the page numbering in this article may contain either page numbering skips, or page numbering repetitions, or both. However, the article content is presented in its entirety and in correct reading order.

Please click on "Next Page" (at the top of the screen) to begin viewing the article.

# [ 127 ]

under which lay those before-mention'd Jaw-bone, and Piece of Horn; which, in all Appearance, to every one that viewed these Stratums, had never been removed.

M. C.

#### Dimensions of the Deers Horns in the Museum of the ROYAL SOCIETY.

		Feet Inches.		
Length of the Skull	-	-	I	4
Breadth of the Forehead		· co-	0	9
Length of each Horn		•	5	0
Distance of the extreme T	ips of the	Horns	6	O

N. B. These Horns (Fig. 2.) are evidently of the same fort as those often found in *Ireland*, of which Descriptions are given in *Transact*.

n. 227, n. 394. and n. 444, p. 389. But I do not remember to have met with any before of this Species found in *England*, or any-where else besides *Ireland*.

C. M

# VI. The Phænomena of Venus, represented in an Orrery made by Mr. James Ferguson, agreeable to the Observations of Seignior Bianchini.

Read March 20. 1745-6. N all the Orreries that I have feen, Venus is represented as having her Axis perpendicular to the Plane of the Ecliptic, and her diurnal Motion thereon equal to 23 Hours of our terrestrial Time. Hence, as her annual Motion is performed in about 225 of our Days, it will contain 234 of hers; consequently, to an Eye placed in Venus, the Sun will always appear to go R

#### T 128 7

thro' a Sign of the Zodiac in  $19\frac{1}{2}$  of her Days; and as her Axis has no Inclination, the must have a continual Equality of her Days and Nights, without any Variation of Scasons, and so her annual Motion can be of no other Use than to keep her from falling down to the Sun.

But Bianchini gives a very different Account of her; which is, that her Axis inclines 75 Degrees from a Line supposed to be drawn perpendicular to the Plane of the Ecliptic (by which I suppose he means her own Ecliptic, and not the Earth's); and that her diurnal Motion is performed in 24 Days and 8 Hours of our Time; and this will cause her Year. which is equal to almost 225 of our Days, to contain only of of her Days; and this odd Quarter of a Day in Venus will make every fourth Year a Leap Year to her, as happens to us on Earth, by the 6 Hours that our Year contains above 365 Days: And to her the Sun will appear always to go thro' a Sign of the Zodiac in little more than \(\frac{3}{4}\) of her Day, which is equal to 181 of our Days; and in going round the Sun, her North Pole constantly leans towards the 20th Degree of Aquarius.

Thus, with regard to the absolute Length of Venus's Year, Bianchini agrees with Cassini and other Astronomers: but differs widely in other very remarkable Particulars, from which arise so many Advantages, as to make that Planet incomparably more sit for its Inhabitants, than we could possibly conceive it to be by a quick Rotation on an Axis perpendicular to its annual Path. For Venus is so much nearer the Sunthan our Earth is, that it is well known she must have twice as much Light and Heat as our Earth has;

# [ 129 ]

and then, was the Sun always perpendicularly above her Equator, we cannot imagine but that her Equatorial Parts must be burnt up with Heat, and her Polar Parts uninhabitable, by reason of the Greatness of Cold, occasioned by the Sun-beams being parallel to, or making so very acute Angles with, the Horizon.

But, by such a Motion as Bianchini describes, and which I have exactly represented in my Orrery, these Inconveniences are avoided; for there is no Place in Venus but what will have the four Seasons every Year, and the heated Places will have Time to cool; because, to any Place over which the Sun passes vertically on any given Day, he will, on the next Day, be 26 Degrees from the Vertex thereof, even tho the Place be on the Tropic; and if it be on the Equator, One Day's Declination will remove him  $37\frac{1}{3}$  Degrees from it.

I having considered in general what the Effects of the Sun's quick and great Declination would be in Venus, as occasioned by the great Inclination of her Axis, with her slow diurnal and quick annual Motion; and finding that her Globe in the Orrery, by being not quite an Inch in Diameter, was insufficient for solving her Phænomena to any Degree of Exactness; I took the following Method, by which I could do it mechanically, to serve my Purpose.

Along the Middle of a strait narrow Slip of Parchment I drew a black Line, and then measuring my Parchment round a common Globe of 9 Inches Diameter, cutting it so as when the Ends were a little overlapp'd, it would become a Girdle, and stick fast on any great Circle of the Globe. Having thus sit-

R 2

ted it, I took it off; and laying it flat on a Table. I divided one Side of the black Line into 97 equal Parts for the o Days and Quarter of a Day in Venus's Year, and then I subdivided each Day into 24 Hours or equal Parts, of which the odd Quarter contained 6, and set the proper Figures to them. The other Side of the Line I divided into 12 equal Parts or Signs, and each Sign into 30 Degrees: By this means I could eafily see, at every Day and Hour in Venus, in what Place of the Ecliptic the Sun was: And putting this Girdle round the Globe, at an Angle of 75 Degrees to the Equator, crosling it in two opposite Points, it would, by representing Venus's Ecliptic drawn on her Globe, serve for the Solution of Problems concerning her, as the Ecliptic on our terrestrial Globe does for those relating to our Earth: For, by bringing the Sun's Place, at any Day or Hour, to the brasen Meridian, I had thereby his Declination for that time; which gave me both an easy and sure Way for drawing the Spiral of the Sun's Motion over the Body of Venus on this Globe; and then, by elevating it to different Latitudes, I could immediately see where the Spirals cut the Horizon in any Latitude, and at what Height or Declination they cross'd the Meridian; as by the Hour-Circle I could easily perceive the Times of the Sun's Rifing and Setting, and his Amplitudes on the Horizon; and I called that the first Meridian. which passed thro' the Northern Tropic, in the Place where the Sun touch'd it at his greatest North Declination; reckoning the East or West Longitudes on the Equator from that Meridian. But this Meridian will only ferve for One Year; because, as the odd Quarter of a Day in Venus causes the Sun to cross

# [ 131 ]

her Equator 90 Degrees Westward of the former Place every Year, the Place of the Sun's greatest Declination at the North Tropic will be in a Meridian 90 Degrees Westward of the former also. Things being thus premis'd in general: I now proceed to give as good a Description as I can of the particular Phanomena in Venus, confining myself chiefly to what happens in her Northern Hemisphere; knowing that the same must happen, mutatis mutandis, in the Southern.

1. Her Axis is inclined  $\int 1\frac{1}{2}$  Degrees more than the Axis of our Earth, and therefore the Variation of her Seasons will be much greater than of ours.

2. Because her North Pole inclines toward Aquarius, and ours to Cancer; her Northern Parts will have Summer in the Signs where those of our Earth have Winter; and vice versa.

3. The artificial Day at each of her Poles (containing  $4\frac{1}{8}$  apparent diurnal Revolutions of the Sun) will be equal to  $112\frac{1}{2}$  natural Days on our Earth.

- 4. The Sun's greatest Declination, on each Side of her Equator, amounts to 75 Degrees: Therefore her Tropics are only 15 Degrees from her Poles, and her Polar Circles at the same Distance from her Equator. Consequently, her Tropics are between her Polar Circles and Poles, contrary to what those on our Earth are.
- 6. The Sun, in one apparent diurnal Revolution from the Equator, and any Meridian where he crosses it, to the same Meridian again, changes his Declination at least 14 Degrees more on Venus, than on our Earth from the Equinox to the Solstice.

#### [ 132 ]

6. Let us now suppose an Inhabitant standing on her North Pole, where the Sun's Declination is al ways the fame with his Altitude, and looking toward that Point of the Horizon where the first Meridian (above-mentioned) cuts it; and let him call that Point the South, so shall he have a Meridian fixt. which will determine the other cardinal Points on the Horizon; tho', strictly speaking, every Point of the Horizon to him is South: Yet, for once, let us suppose him to have an horizontal Plane, fixed with its South Point in this Meridian, and thence divided and numbered like the Horizon of a Globe: Put a moveable Ruler with Sights to turn round the Centre of this Plane, for observing the Sun's Amplitude at Riling and Setting; and a graduated Quadrant to be fixed in the North and South Line, with a moveble Index, for taking the Sun's Altitude, in passing over the Meridian. The fame Degree, or Part of a Degree, that gives him the Altitude, will also give him its Declination, and he will have the following Phanomena.

The Sun will rife  $22\frac{1}{2}$  Degrees North of the East, and going on  $112\frac{1}{2}$  Degrees, as measured on the hozontal Plane, he will cross the Meridian at an Altitude of  $12\frac{1}{2}$  Degrees; then, making an intire Revolution without setting, he will cross it again at an Altitude of  $48\frac{1}{2}$  Degrees: At the next Revolution he will cross it as he culminates, at the Height of 75 Degrees, being only 15 Degrees from the Zenith; and thence he will descend in the like spiral manner, crossing the Meridian first at an Altitude of  $48\frac{1}{2}$  Degrees; then, at an Altitude of  $12\frac{1}{2}$  Degrees, and going on thence  $112\frac{1}{2}$  Degrees he will set  $22\frac{1}{2}$  Degrees North of the West, having been 4 Revolutions and and 5 Parts of One above the Horizon.

# [ 133 ]

- 7. If the Spectator turns his Instrument  $22\frac{1}{2}$  Degrees toward the East, and then supposes his Quadrant in the Plane a new Meridian to him; the Sun will then rise due East, and set in the North-West; and his Declination in the Meridian will not be the same as before; for he will first cross it at an Altitude of 10 Degrees: next of 46; then, of  $74\frac{6}{8}$ ; and, at an Hour and an half after, he will come to his greatest Declination; from which, in his Descent, he will not cross the Meridian in the same Degrees of Altitude, as in ascending he did.
- 8. Now, let the Spectator turn his Instrument 90 Degrees still more toward the East, and the Sun will rise due South; and from thence making a complete Revolution, he will cross the Meridian at an Altitude of 37½ Degrees; making another Revolution, he will cross it at an Altitude of 70¾ Degrees; and, going on 7½ Hours (or 112 Degrees) he comes to his greatest Declination in the West-North-West: Thence descending, at the End of the third Revolution he crosses the Meridian 58½ Degrees high; at the End of the fourth he crosses it in 23¾ Degrees of Altitude; and, going on thence 225 Degrees, or ½ of a Revolution, he sets in the North-East,
- 9. If the Spectator will now turn his Instrument just half round, shifting his Meridian 180 Degrees, the Sun will rise in the North; and, going on 180 Degrees, or half a Revolution, he will cross the Meridian at an Altitude of 19 Degrees; then, making a complete Revolution, he will cross it at an Altitude of 55 Degrees; and, going on thence 292½ Degrees he comes to his greatest Declination in the East-South-East; from which Place he descends, crossing the Meridian in 73½ Degrees of Altitude; and, in the next Revolution, he crosses the

#### [ 134 ]

Meridian at an Altitude of  $41\frac{1}{2}$  Degrees: At the fourth Revolution he crosses it at an Altitude of 5 Degrees; and going on thence 45 Degrees, or  $\frac{1}{8}$  of a Revolution, he sets in the South-West.

10. The Sun being thus for half a Year together above each Pole of Venus in its Turn, will cause the whole Year at her Poles, as well as at the Poles of our Earth, to contain only one Day and one Night: But there, the Difference between the Heat in Summer and Cold in Winter (or of Mid-day and Midnight) is greater than betwixt the same on any two Places of our Earth; because, in Venus, the Sun is for half a Year together above the Horizon of one or other of the Poles; and for at least 3 of a Revolution (or about 16 of our Days) within 20 Degrees of the Zenith! and during the other Half of the Year, always below the Horizon; and for a confiderable Part of that Time, at least 70 Degrees from it: Whereas at the Poles of our Earth, tho' the Sun is for half a Year together above the Horizon, yet his Altitude is never more than 23 Degrees above it in Summer, nor his Depression greater than that Quantity below it in Winter. When the Sun is in the Equator, he is seen in the Horizon of both Poles; one Half of his Difc above, and the other below: And descending quite below the Horizon of one Pole, he afcends in a visible Spiral above that of the other, until he comes within 16 Degrees of the Zenith, where he keeps the same Altitude nearly for some time; then descends in the like spiral manner, till he gets below the Horizon, where he continues invisible for the other Half of the Year. This will occasion to each Pole one Spring, one Harvest, a Summer as long as them both,

#### [ 135 ]

and one Winter, equal in Length to the other three Scasons.

The Sun's great Distance below the Horizon of Venus's Poles, will make her Winters much more uncomfortable than at the Poles of our Earth, where they have Twilight more than half the Winter-time; unless she be surrounded with an Atmosphere capable of occasioning a Twilight, at least as long in proportion to her Winter, as our Twilight is to ours. But this can hardly be suppos'd; because always, when we fee Venus, she appears with the same constant Serenity; and therefore I am apt to beleive she has a Satel lite, to supply, in some measure, the Absence of the Sun; as our Moon does to our Earth's Poles, for one Half of the Winter constantly, without setting, from the first to the third Quarter. 'Tis true, that we are inconveniently polited, with regard to Venus for feeing her Satellite (if she has one); because, when her Moon or Satellite has its enlighten'd Side toward us, it may be too far distant to be seen, because Venus is then beyond the Sun, and, confequently, furthest from us: and when she is betwixt us and the Sun, or thereabouts, her full Moon would have its dark Side to us: And tho' Venus be then nearest the Earth, yet her Satellite could no more be seen by us, than we can see our own Moon at her Conjunction. When Venus is at her greatest Elongation, we should have only one Half of the enlighten'd Side of her full Moon turn'd towards us; and even then, perhaps, on account of its Smallness, it may be too far distant to be seen by our Telescopes. But of this only by-thebye.

11. At the Tropics, the Sun in Summer will continue for about 15 of our Weeks together above the
S
Horizon

# [ 136 ]

Horizon without fetting, and as long below it in Winter without rising. While he is more than 15 Degrees from the Equator, he neither fets to the Inhabitants of the nearest Tropic, nor sets to those of the other; whereas, at our terrestrial Tropics, he rises and fets every Day in the Year. But to let us know more particularly the Phanomena of Venus's Tropics, we will suppose the Inhabitant, who has seen the abovemention'd Appearances at the North Pole, to have travell'd thence along the first Meridian 15 Degrees to the Northern Tropic, carrying his Engine or Infirument along with him; and to have fet it due North and South, in the Place where the faid Meridian interfects the Tropic; and as the Meridian of every Place is in a great Circle passing thro' the Zenith of the Place and both Poles, he can now be at no Loss how to fettle his Meridian, and observe as well the Amplitude and Azimuth, as the Altitude of the Sun; who will rife to him to Degrees North of the East, with about one Degree of North Declination: And going on 100 Degrees (to be measured on the horizontal Plane) he will cross the Meridian with 121 Degrees of North Declination, and 271 of Altitude; then, making an intire Revolution without fetting, he will cross the Meridian at 48½ Degrees of Declination, and 63% of Altitude: At the End of the next Revolution, he will cross the Meridian in the Zenith at the greatest Declination; namely, 75 Degrees; and thence he descends in the like Spiral, crossing the Meridian at the same Altitudes as above, till, in his fifth Revolution, he sets 10 Degrees North of the West.

#### [ 137 ]

12. Let our Traveller now remove Westward on the same Tropic, to a Meridian 97 2 Degrees distant from the first; and there he will have very great Differences of the Rising, Setting, and Meridian Altitude of the Sun; which will now rife to him the first time, in the South Point of his Horizon, at 12 o'Clock; at I o'Clock he will be about half a Degree above the Horizon, and will set at 2 o' Clock: So this short artificial Day in Venus (which is somewhat longer than two natural Days on our Earth) will have no Forenoon at all. The Sun, after continuing almost 14 of Venus's Hours below the Horizon, supposing each diurnal Rotation to be divided into 24 Hours, will rise a little before 4 o'Clock next Morning, near the North-East; and, going on 130 Degrees, he will then cross the Meridian with 22 Degrees of North Declination, and 37 of Altitude: Then, going on without fetting, he again crosses the Meridian at 57 Degrees of Declination, and 72 of Altitude; and advancing forward thence 17 Hours. or 262 Degrees, he comes to his greatest Declination,  $7^{\frac{1}{2}}$  Degrees to the North of the East: From thence, completing his Revolution to the Meridian, he now crosses it in 71½ Degrees of Declination. being only  $3\frac{1}{2}$  Degrees from the Zenith: At the next Revolution he crosses the Meridian with 385 Degrees Declination, and 53½ of Altitude: At the next, which is the fourth Revolution, he croffes the Meridian with 1½ Degree of Declination, and 16½ Degrees of Altitude; and then goes on 65 Degrees, and lets near the West South-West.

13. Suppose now that our Traveller removes still further Westward, on the same Tropic, to a Meri-

2

dian

dian 105 Degrees distant from this his second Station and then the Sun will first rise to him in the South East about 9 o'Clock; and going on thence 45 De grees, he will cross the Meridian with 6 Degrees of South Declination, and 9 of Altitude, at 12: About 2 o'Clock he will be a Degree higher; and, thence descending, he will set near the North West a little before 9 o'Clock: So the Afternoon of this Day is almost 6 Hours (about 6 natural Days with us) longer than the Forenoon; and its Night is but little more than 3 Hours long: For the Sun, after going a little below the Horizon, rises in the North Point thereof; and, making half a Revolution, he croffes the Meridian with 33 Degrees of Declination, and 48 of Altitude; thence, making a whole Revolution, he crosses the Meridian at 66 Degrees of Declination. and 81 of Altitude: At the next Revolution his Declination is 63 Degrees (having passed the greatest 14 Hours before): At the next, it is 28 Degrees of Declination; and, going on thence about 146 Degrees, he fets North West-by-North, about half an Hour after 9 o' Clock; and continues invisible till 3 Quarters past 5 in the next Morning, when he rifes about 4 Degrees North of the East; and, going thence forward 94 Degrees, he crosses the Meridian about 5 Degrees Altitude, and 10 of South Declination, having kept the same Altitude very nearly for three Hours; then descending, he sets in the South South West. about half an Hour past 1 o' Clock; which makes the Afternoon 5 Hours and about 12 Minutes shorter than the Forenoon of the same Day. The Sun now sets for about 15 of our Weeks to Venus's Northern Tropic, and rifes to the Southern; in which the Phanomena

# [ 139 ]

Phanomena are the same: Each Tropic having the four Seasons once every Year; the Winters being longer than the Summers, tho' not quite so long, in proportion, as at the Poles.

14. Having faid so much concerning the North Pole and Tropic, proceed we now to station our Inhabitant in a Place of 45 Degrees of North Latitude, where the first Meridian cuts the Parallel, and he will have the following Phanomena.

The Sun will rife 43 Degrees East of the South, a little before 9 o' Clock; and, ascending very quickly, he will, in little more than 3 Hours, cross the Meridian at an Altitude of 19 Degrees, with 26 Degrees of South Declination; then going on 62 Degrees, he will fet near the West-South-West about 500' Clock in the Afternoon; by which means it is almost two Hours longer than the Fore-noon; each Hour in Venus being equal in Length to 24 Hours and 20 Minutes of our terrestrial Time. The next Day the Sun will rife 3 Degrees North of the East. about half an Hour past 5 o' Clock in the Morning. and will cross the Meridian with 12 T Degrees of North Declination, and 57% of Altitude; and will fet in the North-West-by-West, about half an Hour past 7 o' Clock: So that the Afternoon will be 2 Hours longer than the Forenoon. The next Day the Sun rifes 53 Degrees North of the East, about 3 o' Clock; and will cross the Meridian 31 Degrees North of the Zenith; or with 861 Degrees of North Altitude, and 481 of Declination: Then he goes round without Setting; and crosses the Meridian 30 Degrees North of the Zenith, where he comes to his greatest Declination; from which he returns in the like Spiral toward the Equator,

#### [ 140 ]

Equator, and beyond it; but will not rife and fet at the same Hours as before: For, having made a Revolution without Setting, in the next he fets 53 Degrees North of the West, about o o'Clock: Next Morning he rises in the North-East-by-East, about half an Hour past 4 o' Clock; crosses the Lieridian with 12½ Degrees of Declination, and fets 3 Degrees North of the West, about half an Hour past 6; and now the Forenoon is 2 Hours longer than the Afternoon. The next Day the Sun rises about 7 o' Clock, 62 Degrees East of the South; passes over the Meridian at an Altitude of 19 Degrees, with 26 Degrees of South Declination; and fets a little after 3 o' Clock 3 which makes the Forenoon to be about 2 Hours at least longer than the Afternoon: And now the Sun will continue below the Horizon at least 12 of our Weeks without rising to this Inhabitant of Venus.

15. In this Place of Venus the Hour and Amplitude of the Sun's Rising, for one Half of the Year, are the same with those of his Setting in the other Half; which will also happen in all Places under the first Meridian, where he rises and sets: But, if our Spectator pleases to remove along the Parallel of 45 Degrees Latitude, Eastward 142 Degrees, the Phanomena of Things will then be very different to him; for the Sun once from rifing in the North-East-by-East, will pass over the Meridian with 3 1/2 Degrees of North Declination, and fet due North; which will make the Afternoon somewhat above four Hours longer than the Forenoon; and the next Morning the Sun will rise at 2 o' Clock, 21 - Degrees East of the North, or about the North-North-East. what would happen on the other Days concerning

the

# [ 141 ]

the Sun's Rising and, I shall not take any further Notice of it; but, if the Inhabitant will travel Eastward 371 Degrees, still upon the same Parallel of Latitude, he will see the Sun, at making his first Appearance from the Southern Tropic, rise due South at 12 o' Clock; and. getting about half a Degree above the Horizon, when he has gone forward about o Degrees, he will then descend, and set about a Quarter after 1: So there is only an Hour and a Quarter in the first Day of the Sun's Appearance; and the fecond Day will be 11 Hours long; but the third Day will be about 87 Hours long; for the Sun will make 3 Revolutions and fomewhat more than an half without setting: The fourth Day will II Hours long; and the fifth will only contain an Hour and a Quarter; for the Sun will rife about 18 Degrees East of the South, and set in the South Point of the Horizon.

16. We will now suppose that the Spectator has travelled from 45 Degrees of North Latitude, to the Equator, and has a Mind to take a Tour round the same, because the Phanomena will be very different in different Parts thereof; tho' the Sun will rife and fet to every Part of it, in every apparent Revolution: but we shall only consider in general what happens at two Places thereof: The first Place shall be that, where the first Meridian crosses the Equator; and the second, a Place 112 Degrees Westward of the first. To each of these Places the Sun will always rise at 6. and set at 6, tho' fometimes his Meridian Altitude may be 11 Degrees more or less than his Midnight Depresfion; and in other Places the Difference will amount to 15 or 16 Degrees; so that, if the diurnal and nochurnal

#### [ 142 ]

nocturnal Spirals of the Sun's Motion on the Body of this Planet were measured, the one would very much exceed the other. To the first of these two Places the Sun will rife 74 Degrees South of the East in coming from the Southern Tropic, and fet 611 South of the West, having been 22 Degrees high at Mid-day, and will be 32 depress'd below the Horizon at Midnight. The next Day he will rise 44 Degrees South of the East, and set 26 Degrees South of the West; having been 55 Degrees high at Noon, and will be  $74\frac{1}{2}$  depress'd at Midnight. The third Day he will rise  $7\frac{7}{2}$  Degrees South of the East; and crossing the Equator at half an Hour after 10 o' Clock, he will, in 71 Hours after, fet 12 Degrees North of the West; and so proceed, changing his rising and setting Amplitude every Day, in advancing toward the Northern Tropic, till he reaches it; and then his fetting Amplitude, in going from it, will be the same as his rising Amplitude in coming toward it. fecond Place, all I shall take notice of, is, the Sun, in coming from the Southern to the Northern Tropic, will cross the Equator at 9 o'Clock at Night; and, in going from the Northern to the Southern Tropic, he will cross the Equator at Midday.

17. At the Equator the Sun's Rays will be as oblique, when his Declination is greatest, as they are at London, when he touches the Tropic of Capricorn in December; because the Tropics of Venus are as far from each Side of her Equator, as the Tropic of Capricorn is from the Parallel of London on our Earth: Therefore, at Venus's Equator, there will be two Winters, two Springs, two Summers, and two Autumns, every Year:

# [ 143 ]

Year: And because the Sun stays for some time near the Tropics, and passes so quickly over the Equator, every Winter there will be about twice as long as Summer: But, because of the quick Return of Summers, and the general Heat on the Body of Venus, the Winters there will be very mild; and so will make the Equator, and all Places thereabouts, very temperate, and sit for Habitation.

18. Those Parts of *Venus* which lie between the Poles and Tropics, and between the Tropics and polar Circles, and also between the Polar Circles and Equator, will more or less participate of the *Phanomena* of these Circles, as they are more or less distant from them.

19. The Places of the Equinoxes and Solftices on the Body of Venus go backward, or from East toward the West, 90 Degrees every Year, This is not occasioned by any Mutation of her Axis from its Parallelism; but by the Sun's being a Quarter of a Day later in croffing the Equator every Year, than on the Year before; and therefore he will cross it in a Place 90 Degrees Westward of the former every Year: So that to any Place where he crosses the Equator at Noon, he will, on the Return of that Day at Noon in the next Year, be almost 10 Degrees South of the Equator, and will cross it at 6 in the Evening; supposing the Year to begin when the Sun is on the Equator, in passing from the Southern Tropic to the Northern. Hence, tho' the Spiral, in which the Sun's apparent Motion is performed, be of the same Sort every Year, yet it will not be the very same; because the Sun will pass vertically over all the same Places but once in every four Years: And, in the above Description, I

have

# [ 148 ]

have only shewn what will happen in general, for one Year; having only drawn the Spiral of the Sun's Motion for that Time: And if a Speciator, on any Parallel of Latitude, should want to see the same Appearances of the Sun's Rising and Setting every Year, and, consequently, to have the particular Days thereof to be still of the same Length with those of the Year, he must travel Westward every Year 90 Degrees on the same Parallel.

20. The Inhabitants of Venus will be very careful in adding a Day to some particular Part of every fourth Year, to keep still the same Seasons to the same Times; because, as the great annual Change of the Equinoxes and Solstices will shift the Seasons forward a Quarter of a Day every Year, they would, in 36 Years, shift the Seasons forward thro all the Days of the Year: But, by this intercalary Day, every fourth Year will be a Leap-Year; which will bring her Time to an even Reckoning, and keep her Calendar right.

21. The great Change of the Sun's Declination every Day, which causes his Altitude, at Noon, or any other Hour, and his Amplitude at Rising and Setting, to be so very different in Places lying under the same Parallels of Latitude, will be of one singular Use in Venus, the like whereof we shall never enjoy on the Earth; and that is no less than the giving a sure and casy Method of sinding the Longitude. For, suppose to one Place, at Noon, the Sun's Declination is 30 Degrees, and to another Place it is, only 20 Degrees 35 Minutes at Noon, in the same revolutional Spiral, going from the Equator toward the Northern Tropic; the Difference of these two Declinations is 9 Degrees 25 Minutes: In the same Spiral

# [ 149 ]

Spiral from the Equator, where any Meridian crosses it, to the same Meridian again, the Declination changes from nothing to 37 Degrees 21 Minutes; and the Sun has gone 38 Degrees 55 Minutes in the Ecliptic. These Things being known, the Proportion will be thus; As 75 Degrees, the greatest Declination, is to the Sun's Motion in that Time, which is 3 Signs, equal to 2-5 Revolutions round Venus; so is 9 Degrees 25 Minutes (the Difference of Declination at two given Places) to 9 Degrees 44. Minutes, which is a fourth Part of a Revolution; and therefore the one Place is a fourth Part of a Circle, or 90 Degrees of Longitude distant from the other: And, as the Declination was advancing from the Equator toward the Northern Tropic, the Place, in whose Meridian it was 20 Degrees 35 Minutes, is Eastward from the Place in whose Meridian it was 30 Degrees, supposing them both to be in the Northern Hemifphere.

I should be very glad to see this Description examined into, and put in a better Form, by some whose Abilities are much greater than mine: And altho' it seems strange, at the first View, that the great Inclination of Venus's Axis, with her slow diurnal and quick annual Motion, should make such mighty Differences of her Phanomena from the Earth's; yet I verily believe, that, was the Spiral of the Sun's Motion for sour Years, which would contain 37 Revolutions, nicely drawn on a large Globe, and the Times mentioned in which the Sun would rife and set, with his different Amplitudes, Altitudes, and Declinations, where the Effects thereof would differ considerably in many particular Parts of each Spiral;

 $T_{a}$ 

# [ 146 ]

and so occasion remarkable Differences of the Lengths of Day and Night, in the same Revolutions, to Places under the same Parallels of Latitude; a whole Volume might be wrote in the Description, if the Author would descend to Particulars.

VII. A Machine for founding the Sea at any Depth, or in any Part, invented by Major Wm. Cock in the Year 1738. in a Voyage to Georgia.

Presented April 10. HE Draught of this Machine is exhibited in TAB. II. Fig. 1.

wherein

AAAA represent a Trunk of Timber, with a square Hollow, thro' the Centre of which passes the square Piece of Timber BB.

A Groove on each Side, in which are placed the two Pieces of Iron CC; the Foot of each resting on the Pins DD, that pass thro' the Trunk; the upper Part of the Irons are hooked to an iron Pin at E, which passes thro' the square Piece BB; which Piece is hollowed between H and H, for the Hooks of the Irons CC to pass up and down.

When the Weight F touches the Ground, the two Irons CC fink the Trunk to G, which unhooks them at E; whereupon they fall off, and leave the Trunk at Liberty to float or rise up again to the Surface.

A Machine of these Dimensions, loaded with an iron Ball, F, of 12 Pounds Weight, being let down